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1 Introduction

Interactions between electrons and vibrational modes of solids (phonons) create composite bound states known as polarons. Most of the attention in the field has quite naturally been focused in systems where these effects are expected to be strong, e.g. polar materials,¹ their vacancies,² molecular crystals³ or 2D-materials.⁴ In these systems the electrons interact among others with long-range (LR) dipole fields induced by displaced ions. Although covalent materials have no dipole moments, one could expect long-range quadrupole fields to contribute to the formation of polarons, as inferred from their significant impact on the carrier mobility in Si.^{5,6}

Some covalent systems with strong electron-phonon (e-ph) interactions show conductivity induced by hopping of small polarons, e.g.: disordered systems such as chalcogenide glasses,⁷ molecular crystals such as S₈,⁸ rare gas solids such as Xe,⁹ or 2D phosphorene and arsenene,¹⁰ where the polaron is

Spectroscopic signatures of nonpolarons: the case of diamond*

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Polarons are quasi-particles made from electrons interacting with vibrations in crystal lattices. They derive their name from the strong electron-vibration polar interactions in ionic systems, that induce spectroscopic and optical signatures of such guasi-particles. In this paper, we focus on diamond, a nonpolar crystal with inversion symmetry which nevertheless shows interesting signatures stemming from electron-vibration interactions, better denoted "nonpolaron" signatures in this case. The (non)polaronic effects are produced by short-range crystal fields, while long-range quadrupoles only have a small influence. The corresponding many-body spectral function has a characteristic energy dependence, showing a plateau structure that is similar to but distinct from the satellites observed in the polar Fröhlich case. We determine the temperature-dependent spectral function of diamond by two methods: the standard Dyson-Migdal approach, which calculates electron-phonon interactions within the lowestorder expansion of the self-energy, and the cumulant expansion, which includes higher orders of electron-phonon interactions. The latter corrects the nonpolaron energies and broadening, providing a more realistic spectral function, which we examine in detail for both conduction and valence band edges.

> localized in lone-pair orbitals. Large polarons in covalent materials have been historically neglected, and were dubbed "nonpolarons" by Emin.¹¹ We note that polaronic signatures were found in doped diamond with a hydrogen-terminated surface having a negative electron affinity.¹² In the present paper we study polaronic effects in intrinsic diamond, to find whether the long- or short-range (SR) potential binds phonons to electrons, and to quantify from first-principles the binding and spectral signatures of polarons in non-polar materials.

> The detection of polarons in a crystal often relies on angleresolved photoemission spectroscopy (ARPES), which measures the kinetic energy and angular distribution of electrons excited by incident light. These quantities are directly related to the number of states available, as a function of energy and momentum. Signatures of polarons in ARPES experiments can be found in cuprate superconductors,13 ionic 3D14-17 and 2D15,18 crystals, ferromagnetic materials,19 and interfaces.20 In the simplest model, neglecting surface effects, ARPES can be related to the one-electron spectral function, which is the central property of interest here.

> Besides (non)polaron binding and the renormalization of the direct electronic band gap,²¹ the effects of phonons in nonpolar materials can also be seen in the optical excitation of carriers in indirect band gap semiconductors.²² The scattering of carriers by phonons dominates transport mechanisms at high temperature, hence an appropriate description of the spectral function is essential to calculate transport properties



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(2)

including many-body effects. Important experimental observables are the thermoelectric conductivity,²³ the charge carrier mobility/ conductivity,²⁴ or superconductivity.²⁵

First-principles calculations of the spectral function at the valence band maximum (VBM) and conduction band minimum (CBM) for (polar) LiF and MgO were examined by Nery *et al.*²⁶ within a zero temperature formalism. In this paper, we expand their approach by studying a non-polar material, diamond, and including finite temperature effects. We will focus on the renormalization of electronic energies at T = 0 K (the zero point renormalization, or ZPR), their temperature dependence, and the emergence of nonpolaronic signatures. Vibrational properties and e–ph matrix elements are obtained using density functional perturbation theory (DFPT)^{27,28} while the interaction between electrons and phonons, which leads to the formation of a quasiparticle (QP), is treated using many-body perturbation theory (MBPT).²⁹

The article is organized as follows. Section 2 summarizes the most important theoretical aspects of the e-ph problem with particular emphasis on the different approaches that can be used to compute spectral functions and QP energies. More specifically, we compare the standard Dyson equation in the Migdal approximation (DM)^{30,31} with the cumulant-expansion (CE) method,^{32–34} which includes higher order diagrams in the self-energy. The CE was previously shown to provide accurate results for e-e interactions³⁵ and also to improve plasmonic polaron satellite energies.^{36,37} In Section 3, we describe spectral signatures in diamond and show that the CE gives better results relative to DM. In addition, we show that LR quadrupole fields do not contribute strongly to the spectral signals; the latter are mostly created by local crystal fields. We also include an ESI⁺ which summarizes the basic equations of density functional theory (DFT) (used for ground-state calculations) and DFPT. ESI† also includes analyses and clarifies important aspects of numerical convergence, analytical transformations for the cumulant expansion, and finite temperature effects. Atomic units are used everywhere unless explicitly noted.

2 Methods

We study the interaction between bare electrons and bare phonons by assuming that it can be treated with MBPT techniques. It is important to note that for polarons this is not always possible. In strongly interacting cases, small localized polarons can be formed, which cannot be treated perturbatively.

In ESI[†] S1, we summarize the DFT and DFPT methods. In this section, we detail the MBPT formalism and explain how to split the LR and the SR from the total DFPT potential.

2.1 Self-energy

Following ref. 26, we refer to the lowest order e-ph self-energy (second order in the atomic displacements) as the Fan-Migdal (FM) self-energy.³¹ It includes two terms, the static Debye-Waller (DW)³⁸ and the dynamic Fan³⁹ term,

$$\Sigma_{n\mathbf{k}}(\omega) = \Sigma_{n\mathbf{k}}^{\mathrm{DW}} + \Sigma_{n\mathbf{k}}^{\mathrm{Fan}}(\omega).$$
(1)

where

$$\Sigma_{n\mathbf{k}}^{\mathrm{DW}} = \frac{1}{N_{\mathbf{q}}} \sum_{j\mathbf{q}} g_{nn\mathbf{k}}^{\mathrm{DW},j\mathbf{q},j-\mathbf{q}} \left(2n_{j\mathbf{q}}(T) + 1 \right)$$

and

$$\Sigma_{n\mathbf{k}}^{\text{FAN}}(\omega) = \frac{1}{N_{\mathbf{q}}} \sum_{m} \sum_{j\mathbf{q}} |g_{nm\mathbf{k}}^{j\mathbf{q}}|^{2} \\ \times \left(\underbrace{\frac{n_{j\mathbf{q}}(T) + f_{m\mathbf{k}+\mathbf{q}}(T)}{\omega - \varepsilon_{m\mathbf{k}+\mathbf{q}} + \omega_{j\mathbf{q}} + i\eta}}_{\Sigma^{-}} + \underbrace{\frac{n_{j\mathbf{q}}(T) + 1 - f_{m\mathbf{k}+\mathbf{q}}(T)}{\omega - \varepsilon_{m\mathbf{k}+\mathbf{q}} - \omega_{j\mathbf{q}} + i\eta}}_{\Sigma^{+}} \right)$$
(3)

with η a positive infinitesimal (retarded self-energy).⁴⁰ In the equations above, we use $\varepsilon_{m\mathbf{k}}$ to denote the energy of the electronic state with band index *m*, wavevector \mathbf{k} , and Fermi-Dirac occupation number $f_{m\mathbf{k}}$. The phonon frequencies are denoted as $\omega_{j\mathbf{q}}$, with *j* the mode index, \mathbf{q} the phonon wave vector, and $n_{j\mathbf{q}}$ the Bose–Einstein occupation number. Finally, the symbol $g_{nm\mathbf{k}}^{j\mathbf{q}}$ denotes the first order e–ph matrix element.

In principle, the DW matrix element g^{DW} should be computed as the second-order derivative of the self-consistent potential with respect to the phonon displacement. In practice, we use the acoustic sum rule, together with the rigid-ion approximation,⁴¹ to express g^{DW} in terms of the first order e-ph matrix elements.^{40,42} Inside the parentheses of eqn (3), there are two terms with the Bose Einstein factors *n* and *n* + 1, which we label Σ^- and Σ^+ . Each term represents in turn two separate scattering events, into and out of state *n***k**, which correspond to absorption and emission of phonons, pairing with the appropriate electron or hole state.

2.2 Dyson-Migdal approach

The interacting Green's function *G* can be expressed in terms of the initial bare propagator G^0 and the exact self-energy *via* the Dyson equation.⁴³ In the diagonal approximation,⁴⁴ the off-diagonal matrix-elements of Σ in the Bloch basis set are assumed to be negligible, and the Dyson equation reduces to

$$G_{n\mathbf{k}}(\omega) = \frac{1}{\left(G_{n\mathbf{k}}^{0}(\omega)\right)^{-1} - \Sigma_{n\mathbf{k}}(\omega)}.$$
(4)

Finally, the spectral function A is given by

$$A_{n\mathbf{k}}(\omega) = -\frac{1}{\pi} \Im m G_{n\mathbf{k}}^{\mathbf{R}}(\omega).$$
(5)

where G^{R} is the retarded Green's function. From eqn (4) and (5), one obtains⁴⁰

$$A_{n\mathbf{k}}(\omega) = -\frac{1}{\pi} \frac{\Im \Sigma_{n\mathbf{k}}(\omega)}{(\omega - \varepsilon_{n\mathbf{k}} - \Re \Sigma_{n\mathbf{k}}(\omega))^2 + \Im \Sigma_{n\mathbf{k}}(\omega)^2}$$
(6)

At this point, it is worth stressing that in practical applications it is customary to evaluate eqn (6) using the lowest order FM self-energy, eqn (1), evaluated with bare electron/phonon quantities (one-shot method). This approach, which neglects self-consistency effects and vertex corrections,^{26,40} will be referred

Paper

to as the Dyson–Migdal (DM) approximation in what follows. According to previous studies in polar materials,²⁶ the DM approach usually yields poor QP energies and spectral weights when compared with high-quality Monte Carlo calculations for the Fröhlich model.⁴⁵ Moreover, the position of the DM satellite relative to the QP peak is often inaccurate and far from the expected value, which should match the phonon frequency $\omega_{\rm LO}$ of the LO mode. A promising route for going beyond the DM approximation is the cumulant expansion detailed in the next section.

2.3 Cumulant expansion

The Green's function in the time domain can be rewritten in an exponential form (cumulant expansion) using Kubo's formula³²

$$G_{n\mathbf{k}}(t) = G_{n\mathbf{k}}^0(t) \mathrm{e}^{C_{n\mathbf{k}}(t)},\tag{7}$$

where

$$C_{n\mathbf{k}}(t) = \sum_{i=2}^{\infty} C_{n\mathbf{k}}^{(i)}(t)$$
(8)

is the sum of the cumulants of the *i*-th order. The cumulant function in eqn (8) can achieve accurate results⁴⁶ for our problem using just i = 2, and is exact for a (fully localized) core electron interacting with a phonon.⁴⁷

The cumulant functions can be determined by expanding the exponential in eqn (7) and comparing powers with the standard Feynman expansion of the Green's function.³⁵ Using the Fan–Migdal self-energy (Fan and DW terms), one obtains

$$C_{n\mathbf{k}}(t) = \int \mathrm{d}\omega\beta_{n\mathbf{k}}(\omega) \frac{\mathrm{e}^{-i\omega t} + i\omega t - 1}{\omega^2}$$
(9)

where

$$\beta_{n\mathbf{k}}(\omega) = \frac{1}{\pi} |\Im m \Sigma_{n\mathbf{k}}^{\mathrm{Fan}}(\omega + \varepsilon_{n\mathbf{k}})|$$
(10)

while the DW self-energy appears as a pure shift of the QP energy,

$$G_{n\mathbf{k}}(t) = -i\theta(t)\mathrm{e}^{-i\left(\varepsilon_{n\mathbf{k}} + \Sigma_{n\mathbf{k}}^{\mathrm{DW}}\right)t}\mathrm{e}^{C_{n\mathbf{k}}(t)}.$$
 (11)

The three terms in eqn (9) have different effects on the spectral function. The first one gives rise to satellites, the second term shifts the QP peak, while the third term corrects the QP weight. The second term is calculated using the Kramers–Kronig relations. Further details can be found in Section S3 of the ESI.[†]

The spectral function is obtained by applying the Fourier transform to the Green's Function in the time domain, eqn (11), and inserting it into eqn (5). It can be shown that the CE Green's function is exact to second order, and all higher order terms are included, though in an approximate way, while DM only includes the exact second order terms.^{48,49} In the long time limit, the cumulant has an affine asymptotic behaviour which contains its contributions to the lifetime and lineshape of the QP state,

$$G_{n\mathbf{k}}(t \to \infty) \approx G_{n\mathbf{k}}^0(t) \exp[-w_k + (-\Gamma_{n\mathbf{k}} + i\Lambda_{n\mathbf{k}})t]$$
(12)

where w_k is a constant,⁴⁶ $\Gamma_{n\mathbf{k}} = |\Im m \Sigma_{n\mathbf{k}}(\omega = \varepsilon_{n\mathbf{k}}^0)|$ is the decay rate, $\Lambda_{n\mathbf{k}} = \Re e \Sigma_{n\mathbf{k}}(\omega = \varepsilon_{n\mathbf{k}}^0)$, and G^0 depends on *t* through an exponential with the bare electron energy.

2.4 Energy renormalization

In this section, we summarize three commonly used approximations to compute the QP energy from the e-ph self-energy. If we ignore the frequency dependence of the imaginary part of the self-energy, one obtains that the main peak of the DM spectral function (eqn (6)) is located at the energy $\varepsilon_{n\mathbf{k}}^{\text{NL}}$ that solves the non-linear (NL) QP equation

$$\varepsilon_{n\mathbf{k}}^{\mathrm{NL}} = \varepsilon_{n\mathbf{k}}^{0} + \Re e \Sigma_{n\mathbf{k}} (\omega = \varepsilon_{n\mathbf{k}}^{\mathrm{NL}}).$$
(13)

This equation must be solved numerically using, *e.g.*, rootfinding algorithms that require the knowledge of $\Sigma_{n\mathbf{k}}(\omega)$ for several frequencies. The problem can be significantly simplified if we assume the QP correction $\varepsilon_{n\mathbf{k}}^{\mathrm{NL}} - \varepsilon_{n\mathbf{k}}^{0}$ to be small. In this case, one can expand the self-energy to linear order around the KS energy $\varepsilon_{n\mathbf{k}}^{0}$ to obtain the linearized QP equation

$$\varepsilon_{n\mathbf{k}}^{\text{linear}} = \varepsilon_{n\mathbf{k}}^{0} + Z_{n\mathbf{k}} \Re e \Sigma_{n\mathbf{k}} (\omega = \varepsilon_{n\mathbf{k}}^{0})$$
(14)

with the renormalization factor $Z_{n\mathbf{k}}$ given by

$$Z_{n\mathbf{k}} = \left(1 - \Re e \frac{\partial \Sigma_{n\mathbf{k}}(\omega, T)}{\partial \omega} \Big|_{\omega = \varepsilon_{n\mathbf{k}}^{0}}\right)^{-1}, \quad (15)$$

that is approximately equal to the area under the QP peak of the spectral function. Finally, in Rayleigh–Schrödinger perturbation theory, also known as the on-the-mass-shell (OMS) approach, the energy correction is just given by the self-energy evaluated at the KS energy⁴⁰

$$\varepsilon_{n\mathbf{k}}^{\text{OMS}} = \varepsilon_{n\mathbf{k}}^{0} + \Re e \Sigma_{n\mathbf{k}} \left(\omega = \varepsilon_{n\mathbf{k}}^{0} \right). \tag{16}$$

The evaluation of the energy correction is also carried out for the CE approach. The effective CE self-energy is determined by inverting eqn (4) after inserting the CE and non-interacting Green's functions. Then the NL equation is used to obtain eigenenergies. Further details over the effective CE self-energy are described in S5 (ESI[†]).

In Section 3.1, we will check the results of the different approaches for the band edges of diamond, showing that the OMS approximation using DM produces QP shifts very close to the CE-NL (as already observed for polar materials²⁶).

2.5 Long-range and short-range potentials

Converging e-ph calculations requires very dense **q**-grids that are prohibitively expensive for DFPT. For this reason, we employ the Fourier-based interpolation scheme, initially proposed by Eiguren,⁵⁰ to interpolate the e-ph scattering potentials on arbitrarily dense **q**-meshes. The potential interpolation scheme has an important advantage for the e-ph matrix elements of high unoccupied states, which are very delicate to access accurately with Wannier Function based interpolation methods.⁵¹

The e-ph scattering potential presents a non-analytical behaviour⁵² in the long wave-length limit ($\mathbf{q} \rightarrow 0$) associated to a LR behaviour in real-space, which requires a specialized numerical treatment. Over the past decade, this problem has been subject to several investigations that lead to a well established procedure: the dipole fields for $\mathbf{q} \rightarrow 0$ are treated

using a Fröhlich-like potential,^{53,54} which depends on the Born effective charges and diverges as 1/*q*. Recently,^{5,55,56} the treatment of LR contributions has been generalized to include contributions generated by dynamical quadrupoles.⁵⁷ As the Born effective charges of diamond are zero, in what follows we focus on the treatment of the quadrupole interaction.

In non-polar materials, the Fourier interpolation of the e–ph potentials proceeds by first removing the non-analytical long-range contribution induced by the displacement of the κ -th atom along the Cartesian direction α using the generalized quadrupole model

$$V_{\kappa\alpha,\mathbf{q}}^{\mathscr{Q}}(\mathbf{r}) = \frac{4\pi}{\Omega} \sum_{\mathbf{G}\neq-\mathbf{q}} \frac{(q_{\beta}+G_{\beta})(q_{\gamma}+G_{\gamma})}{(q_{\delta}+G_{\delta})\varepsilon_{\delta\delta'}^{\infty}(q_{\delta'}+G_{\delta'})} \times \frac{1}{2} \mathcal{Q}_{\kappa\alpha}^{\beta\gamma} e^{i(q_{\eta}+G_{\eta})(r_{\eta}-\tau_{\kappa\eta})} e^{-\frac{|\mathbf{q}+\mathbf{G}|^{2}}{4\alpha}},$$
(17)

where Ω is the unit cell volume, Q is the dynamical quadrupole tensor, G are the reciprocal lattice vectors, τ_{κ} is the position of the atom in the unit cell, and ϵ^{∞} is the electronic dielectric tensor in Cartesian coordinates (repeated indices are implicitly

summed over). The last exponential term, $e^{-\frac{|\mathbf{q} + \mathbf{G}|^2}{4\alpha}}$, is a Gaussian filter⁵⁵ with a variance of $\sqrt{\alpha}$. The resulting short-range potentials, which are smooth and analytic in \mathbf{q} space, are then used to build the scattering potential $W_{\kappa,\alpha}(\mathbf{r}, \mathbf{R})$ in the real-space supercell (see, *e.g.*, eqn (12) in Brunin *et al.*⁵⁵). The short-range *W* is Fourier-interpolated on a much denser \mathbf{q} -grid and the non-analytic LR terms are finally added back to get the total scattering potential. Interestingly, one can use this interpolation technique to compute e-ph matrix elements in which only the LR (SR) part of the scattering potential is included. In Section 3.2, we will analyze the separate contributions from the SR and LR potential and their effect on self energies and spectral functions.

3 Results

All calculations in this work are performed using the ABINIT^{58,59} software. Convergence studies are presented in this section, followed by an analysis of spectral functions and ARPES spectra. All calculations use the full DFPT potential, albeit at the end of this section, we split the scattering potential into LR and SR contributions and check their respective effects on the QP. Further details on ground-state and DFPT calculations, such as the lattice constant, band gap and phonon band structure, can be found in ESI† S2.

3.0 Convergence studies

3.0.1 Zero-point renormalization. Accurate ZPR calculations require a careful convergence study with respect to the BZ sampling and the finite value of η . Fig. 1 shows the results of such a convergence study for the CBM of diamond. A uniform $16 \times 16 \times 16$ q-grid sampling with $\eta = 10$ meV gives converged values for the ZPR, but we will see that other quantities are more sensitive. Throughout the paper, we will refer to a **q**-grid

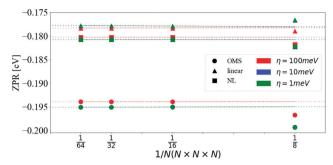


Fig. 1 ZPR for the CBM as a function of the number of divisions in the **q**-mesh and different values of η . The data was fitted with a linear model and extrapolated for $1/N \rightarrow 0$. Calculations were done with the OMS approach eqn (16) (circles), the linear approximation eqn (14) (triangles) and the non-linear approach eqn (13) (squares). The values obtained with $\eta = 10$ meV and $\eta = 1$ meV are indistinguishable on the scale of the graph.

of size $N \times N \times N$ using just *N*. The ZPR of the VBM converges with the same parameters as that of the CBM, giving a total ZPR for the band gap of -0.325 eV within the OMS equation. The ZPRs of the VBM, CBM, and the band gap obtained with the three approximations, discussed in Section 2.4, are summarized in Table 1. At this level of theory, we expect OMS to provide the most accurate results by analogy with polar materials, where OMS is in good agreement with high-quality Monte Carlo methods.²⁶ This might occur due to a fortuitous cancellation between the errors coming from the lack of higher order e–ph interactions, and the evaluation of the self-energy at the KS energy, a non-self-consistent calculation of the QP energy.

The ZPR of polar materials, such as LiF and MgO, is largely dominated by the Fröhlich interaction.²⁶ In diamond, there are no dipole contributions, yet the ZPR is similar to that of LiF and MgO in relative terms. The Fröhlich ZPR of $MgO^{26,60}$ and $LiF^{26,61}$ for the CBM is approximately 4% and 1%, respectively, of the experimental band gap. Similarly, the ZPR of diamond⁶² is about 4%. This shows that accurate computations of band gaps require the inclusion of e-ph interaction even in homo-polar crystals. It should be noted, however, that electron-electron interactions beyond the KS-DFT mean-field approximation are absent in our calculations. These corrections may vary depending on the wavevector, e.g., for the indirect band gap of diamond the GW correction is around 0.02 eV while it is about 0.2 eV for the direct band gap.^{21,63} We also ignored thermal expansion, zeropoint lattice expansion,^{64,65} and further anharmonic effects. Other calculations including these phenomena are detailed in Table 2.

3.0.2 Interplay between $\Im m\Sigma$ and η . In this subsection, we analyze the convergence of the imaginary part of the e-ph

Table 1 Converged ZPR values [eV] for CBM, VBM and band gap at T = 0 Kusing the OMS approach eqn (16), the linear approximation eqn (14)) andthe non-linear approach eqn (13)

	CBM	VBM	Band gap	
OMS	-0.196	0.130	-0.325	
Linear	-0.177	0.118	-0.295	
NL	-0.180	0.119	-0.299	

Paper

Table 2 Diamond indirect band-gap ZPR calculations found in the literature. Meaning of the abbreviations: All calculations were done within firstprinciples (FP), except the last one * which used a semi-empirical exchange-correlation functional (Semi-Emp.); Allen, Heine and Cardona approach (AHC); Supercell (SC) calculations; thermal expansion of the crystal (TE); calculations with electron–electron (EE) interactions; frozen-phonons approximation (FP); harmonic approximation (HA); anharmonic effects (AnHA); (?) not explicitly clarified in the references, but presumably harmonic

ZPR (meV)	Calc./exp.	DF(P)T/MC	SC/AHC	(Non-)A	EE	TE	(An)Ha
2 5 170	-			. ,			. ,
-364^{70}							
-325	FP (This work)	DFPT	AHC	Non-A	No	No	На
-315^{63}	FD	MC	SC	А	No	Yes	2
-320^{44}		DFT	\mathbf{SC}	Α	No	NO	AnHa
-330^{42}	FP	DFPT	AHC	Non-A	No	No	Ha
-337^{63}	FP	MC	SC	Α	G_0W_0	Yes	?
-366^{44}		DFPT	AHC	Non-A	No	No	На
-372^{44}		DFPT	AHC	Α	No	No	На
-380^{42}		DFPT	AHC	Α	No	No	На
-436^{44}		DFPT	SC	Α	No	No	На
-439^{44}	FP	DFT	SC	Α	No	No	На
-462^{71}	FP	DFT	SC	Α	No	No	На
-619^{72}	Semi-Emp*	DFPT	AHC	Α	No	Yes	На

self-energy with respect to the **q**-sampling and the broadening parameter η . This convergence study is needed because the CE is rather sensitive to the quality of the input FM self-energy as detailed in the next section. According to our numerical tests, indeed, the real part of the self-energy at the KS energy converges with relatively coarse **q**-meshes and large η provided that enough empty states are included in the calculation. On the contrary, the imaginay part requires much denser **q**-grids and smaller η . To elucidate this point, we compare finite- η results with those obtained with the more accurate linear tetrahedron method⁶⁶ that we considere as a reference value.

Fig. 2 shows the convergence of the imaginary part of Σ at the CBM using the tetrahedron scheme. Above 5 K, $\Im m \Sigma(\varepsilon_{\text{CBM}}^0)$ converges at N = 160, and from 5 K to 60 K, there is a steep increase of $\Im m \Sigma(\varepsilon_{\text{CBM}}^0)$ from 6×10^{-8} eV to 10^{-5} eV, respectively. At 300 K, $\Im m \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0)$ reaches 10^{-4} eV and at 1500 K it is almost 10^{-2} eV. In Fig. 3 and 4, we compare finite η results with the converged tetrahedron values.

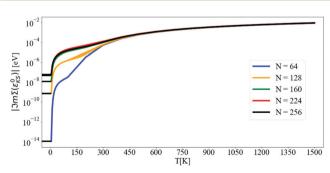


Fig. 2 Convergence with *N* of the imaginary part of the self-energy at the CBM evaluated at the KS energy as a function of temperature. Calculations are done with the DM approach and the tetrahedron method. The lowest *T* considered is 5 K. Each curve has been extended to negative *T* to better visualize the convergence. The *y*-axis is in logarithmic scale.

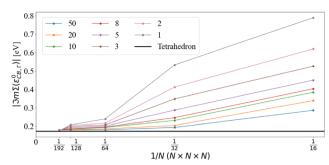


Fig. 3 Convergence of $\Im m \Sigma_{cr}(\varepsilon_{cr}^0)$ using the DM approach with *N* for different values of $\eta = 50, 20, 10, 8, 5, 3, 2$, and 1 meV (indicated in the legend) at 300 K. The black line denotes the reference value obtained with the tetrahedron method and N = 160.

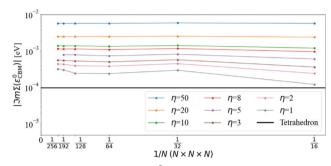


Fig. 4 Convergence of $\Im m \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0)$ using the DM approach with *N*, for different values of η , at 300 K. Exactly at the CBM, the standard method converges slowly down to the tetrahedron value, as the value of Σ is very small and comparable to the "infinitesimal" η itself. The tetrahedron value is calculated with N = 160. The *y*-axis is in logarithmic scale.

Fig. 3 shows that $\Im m \Sigma_{c\Gamma}(\varepsilon_{c\Gamma}^0)$, where $c\Gamma$ means bottom of the conduction band at Γ , converges towards the tetrahedron value for large *N* and small η . Using N = 128, $\Im m \Sigma_{c\Gamma}(\varepsilon_{c\Gamma}^0)$ differs from the tetrahedron method by values in the 34 meV to 6 meV interval when varying η from 50 meV to 1 meV, respectively. Increasing the **q**-mesh density to N = 192, the interval is even smaller, going from 6 meV to 1 meV when varying η from 50 meV to 1 meV.

The convergence of the imaginary part at the CBM is more problematic (see Fig. 4): the value of $\Im m \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0)$ at T = 300 K with the tetrahedron method is very small, around 10^{-4} eV, *i.e.* smaller than the values of η . Further convergence would require $\eta \leq 10^{-4}$ and even denser grids, which are not practical or indeed necessary. Selecting $\eta = 5$ meV and N = 128 (values that will be used later on), $\Im m \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0)$ is within an order of magnitude of the very small tetrahedron value. Lowering η systematically decreases $\Im m \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0)$, but it also increases numerical noise as detailed below.

The effect of increasing the density of the **q**-mesh in the imaginary part of the dynamical self-energy can be observed in Fig. 5: as the **q**-mesh increases, the noise decreases. We choose a **q**-mesh of N = 128 for all the following calculations, which produces low computational noise and a convergence error of 0.6% on the value of the self energy at the CBM.

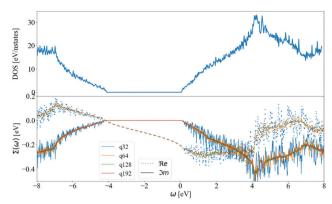


Fig. 5 The (a) density of states and (b) self-energy at CBM and 300 K calculated using DM and $\eta = 5$ meV for N = 32, 64, 128 and 192 at CBM. The DOS is plotted as a reference, showing its similarity with the imaginary part of the self-energy. Self-energy becomes smooth and converged at N = 128 in green (difficult to visualize) with very small oscillations around the N = 192 data in red (easy to visualize).

In Fig. 5, we also display the density of states (DOS). The imaginary part of the self-energy can be seen as a similar sum over accessible electronic states at m, $\mathbf{k} + \mathbf{q}$, but weighted by finite e-ph matrix elements, Bose–Einstein factors at finite T, and shifted by phonon frequencies. Another observation is a linear departure of $\Im m \Sigma_{\text{CBM}}(\omega)$ at ω_{LO} from the value $\Im m \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0)$. The DOS shows a similar behavior. This is at variance with the case of Fröhlich model and polar materials,²⁶ in which the e-ph matrix elements diverge at small \mathbf{q} . This results in a peak in the self-energy at the LO phonon frequency, which is not present in the corresponding DOS.

For a fixed, accurate, **q**-mesh size (N = 128), we show the effect of η in Fig. 6. At very low $\eta = 1$ meV, there are strong oscillations in the satellite features of the spectral function. The width of the QP peak also increases with η , such that a numerical compromise must be found. Above the CBM QP energy of -0.180 eV, there are two plateaus in the spectral function: a small one starting at zero frequency, and a more important one starting at the highest phonon frequency, ω_{LO} , (dot-dashed line), analogous to the peak observed in the Frölich

self energy form for polar materials. To avoid computational noise, and very dense **q**-meshes, we select for the rest of the paper a N = 128 **q**-mesh and $\eta = 5$ meV. Analysis of the plateaus and the spectral function will be detailed in Section 3.1.

Other convergence to consider in the self-energy is the sum over *m* bands in eqn (3), which includes unoccupied states and its numerical convergence over empty bands becomes burdensome. An alternative is to replace high-energy bands with the solution of a non-self-consistent Sternheimer equation.⁶⁷ Study of this convergence can be found in Section S4 (ESI[†]).

The calculation of the cumulant function at low temperatures is complicated by the very small value of $\Im m \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0)$. $\Im m \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0)$ is related to the decay of the QP and its lifetime: the smaller it is, the longer it takes the QP to decay. The range of the time mesh can be estimated through eqn (12), which is an envelope of the Green's function, which can be used to determine the time at which the QP decays and reaches a given tolerance tol, $t_{\text{max}} = -(\ln(\text{tol}) + w)/\Gamma$.

At low temperatures, *e.g.* 40 K, $\Im m \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0)$ is -4.95×10^{-6} eV and the number of frequency points required is huge. The grid size can be estimated as $\Delta \omega / \delta \omega$, where $\Delta \omega$ is the frequency range, and $\delta \omega$ is the spacing between points, which can be otained as $\delta \omega = 2\pi / t_{\text{max}}$. We get $t_{\text{max}} \approx 918$ ps for a tol of 10^{-3} , and for a range $\Delta \omega$ of 20 eV (the convergence of $\Delta \omega$ is explained in S3, ESI†), we obtain 5.3×10^6 points. Therefore, we limit the calculations to room temperature and above, where considerably fewer frequency points are required.

At T = 300 K, $\Im m \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0) = -1.51 \times 10^{-4} \text{ eV}$ gives a $t_{\text{max}} \approx$ 20 ps with a tolerance of 10^{-3} , leading to 174.7×10^3 frequency points. Increasing the temperature, $\Im m \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0)$ increases, which means shorter QP lifetimes and a lower number of frequency points.

In Fig. 7, the cumulant function is shown for T = 300, 900, and 1500 K. The linear behavior of the cumulant at infinite time is present both in the real and imaginary parts. The large time slope of the real part is $\Im m \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0)$ and the slope of the imaginary part is $\Re e \Sigma_{\text{CBM}}(\varepsilon_{\text{CBM}}^0)$. The intercept of $\Re eC(t)$, w_k , originates from the asymmetry⁴⁶ of the denominator of Σ when summing over all $j\mathbf{q}$ states, both elements described in eqn (12) and its convergence is examined in more detail in the S3 (ESI[†]).

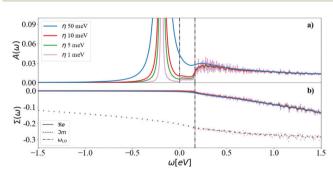


Fig. 6 The (a) spectral function and (b) self-energy at the CBM and 300 K calculated using DM and a N = 128 for different values of $\eta = 50, 10, 5$ and 1 meV. The first vertical line is at 0 energy (CBM) and the second at the highest phonon frequency. The finite temperature allows for absorption of acoustic phonons, which produces a small plateau between the QP peak and the main (optical) satellite feature.

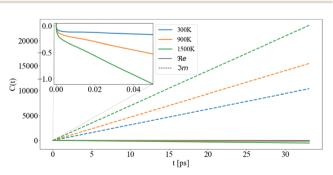


Fig. 7 The cumulant function for CBM between 0 and 32 ps, for different temperatures: 300, 900, and 1500 K. The solid and the dashed lines are the real and imaginary parts of the cumulant function, respectively. The inset shows the non-linear behaviour of $\Re eC(t)$ at small *t*.

The increase of the $\Re eC(t)$ at $t \to \infty$ accelerates the decay of the QP and determines its lifetime $\tau = 1/2\Gamma$.

3.1 Spectral function and ARPES at finite temperature

After converging the self-energy, we can evaluate the spectral function at finite temperatures using the DM and the CE approaches. In Fig. 8, the ZPR can be observed by following the shift of the renormalized band gap to T = 0 K (the bare band gap is set to zero). The ZPR can be determined experimentally by extrapolating the linear regime at high temperatures to 0 K. Using a Pässler fit⁶⁸ to the measurements of Clark *et al.*⁶⁹ gives a ZPR of -0.259 eV. Using also the difference in renormalization for isotopes ¹²C and ¹³C, the obtained ZPR is -0.364 eV.⁷⁰ Table 2 shows a range of calculated ZPR that go from -315 meV to -619 meV.

The fast convergence of the real part of the self-energy allows to determine the renormalization of the band gap in Fig. 8 at a coarse N = 64 **q**-mesh and at a η of 10 meV. The slope of the DM-OMS energy is very close to the CE-NL slope, $-0.409 \text{ meV K}^{-1}$, giving an extrapolated ZPR from high temperatures of -0.281 eV. This value differs from the CE-NL calculated at T = 0 K by -0.066 eV. This discrepancy emerges from two factors: the linear extrapolation should be made above diamond's high Debye temperature (2246 K),73 and the non-adiabatic self-energy, eqn (3), includes a plus and minus term of the phonon energy that is not present in the adiabatic self-energy (an adiabatic selfenergy yields the same ZPR when calculated at T = 0 K and when extrapolated from high temperatures). The very similar behaviour with temperature for DM-OMS and CE-NL derives from the cancellation of errors of the former between the exclusion of higher order e-ph interactions and the evaluation of the selfenergy at the KS energy. The behavior of DM-NL is erratic, since it is not linear at high temperatures, as opposed to CE-NL and DM-OMS. If one attemps a linear extrapolation for temperatures above 2500 K, where it becomes almost linear, the ZPR gives -0.092 eV with a discrepancy of -0.207 eV from the ZPR calculated at T = 0 K, three fold the CE-NL discrepancy. Note that

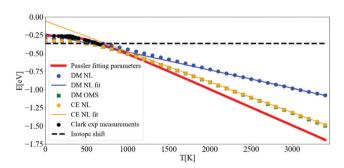


Fig. 8 Variation of the indirect band gap of diamond with respect to temperature. The bare band gap was set to zero. The calculations of the band gap using DM-NL (blue points) do not produce a strictly linear behaviour, and the DM-OMS (green squares) produces the same values as the CE using the NL approach (yellow points). Pässler fit⁶⁸ (solid red line) to the experimental points by Clark *et al.*⁶⁹ (black points) and measured ZPR by isotope shift⁷⁰ (black dashed line). The calculations were done with N = 64 and $\eta = 10$ meV.

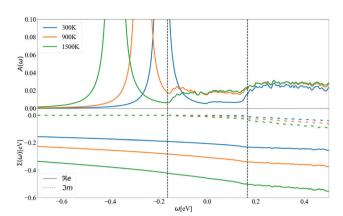


Fig. 9 (a) The DM spectral function and (b) the real and imaginary parts of the self-energy at the CBM for *T* = 300, 900 and 1500 K, with *N* = 128 and η = 5 meV. The vertical dashed black lines correspond to the phonon energies $-\omega_{\rm LO}$ and $\omega_{\rm LO}$ around the KS energy that is used as origin of the *x*-axis.

in the NL approach we use the value of the highest point of the main peak in the spectral function (as it will be seen later on) and should be the one taken into account. The fortuitous coincidence with DM-OMS has been found in other works.²⁶

Going further by evaluating the dynamical part of the spectral function at the CBM in Fig. 9 for T = 300, 900, and 1500 K, one can observe a temperature-dependent offset in the QP peak (the KS energy is $\omega = 0$) which corresponds to the DM-NL energy. Unlike the Fröhlich polaron in simple polar materials, where there are well defined satellite peaks, here the signature of a nonpolaron is a plateau. As the temperature increases, more states become available; in particular, Σ^- starts to contribute to the presence of another plateau with states below the KS energy. Similar to polar materials, the DM approach overestimates the energy distance between the main QP and the high plateau, giving 2.4 ω_{LO} at T = 300 K.

The energy shift of the QP relative to the KS energy at both the CBM and VBM reduces the band gap, as shown in Fig. 10 and 11,

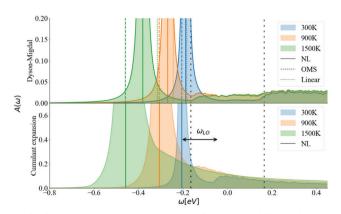


Fig. 10 Spectral function calculated at the CBM at 300, 900, 1500 K using DM and CE methods. The vertical colored lines show the QP energy calculated using non-linear (NL), on-the-mass-shell (OMS) and linear approaches. The vertical dashed black lines correspond to $\pm \omega_{LO}$. η is set to 5 meV and N = 128.

PCCP

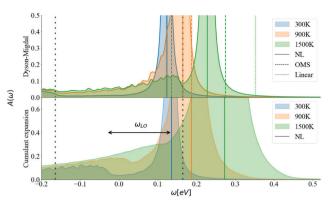


Fig. 11 Spectral function calculated at the VBM at 300, 900, 1500 (K) using DM and CE methods. The vertical colored lines show the QP energy calculated using non-linear (NL), on-the-mass-shell (OMS) and linear approaches. The vertical dashed black lines correspond to $\pm \omega_{\rm LO}$. η is set to 5 meV and N = 128.

respectively (the KS energy is set at 0 eV in each case). Taken together, they determine the band gap shift in Fig. 8. The QP peaks are located at the $\varepsilon^{\rm NL}$ energy. The DM-NL energy differs from the CE-NL more noticeably at higher temperatures. Since the latter contains higher orders of e-ph interactions, the CE approach is more appropriate to determine the QP energy. These energies are very close to the DM-OMS approximation, as also observed in Fig. 8.

At the DM level, the frequencies of both plateaus are disconnected from the QP energy, as discussed in the ESI⁺ S6.2. As the temperature increases, the nonpolaron plateau seems to get closer to the QP peak. However, this is an artifact of DM, as new states become accessible through Σ^- . In DM, Σ^+ gives contributions that start at $+\omega_{LO}$ counting from the KS energy, independently of the temperature or position of the QP peak. At higher temperatures, Σ^- becomes larger and contributes to the spectral function at $-\omega_{LO}$. Since the band renormalization is larger than $-\omega_{\rm LO}$, the $-\omega_{\rm LO}$ plateau still appears to the right of the QP peak, giving the impression that the plateau is shifting with temperature, while it actually corresponds to another plateau. Also, there is an un-intuitive behaviour with DM at the VBM, where the broadening of the QP at 900 K seems wider than at 1500 K. This is due to the overlap between the states created by Σ^+ at $+\omega_{LO}$ (counting from $\omega = 0$) and the QP peak. More detailed analysis is deferred to S6.2 (ESI⁺).

For the CE approach, let us focus first on the CBM. At low temperatures, Σ^+ produces a plateau at the right energy distance, $+\omega_{\rm LO}$, from the QP peak. When increasing the temperature, the main QP becomes wider. It is easier to identify features in the DM spectral function, because the self-energy consists only of two terms, Σ^+ and Σ^- . In the CE instead, as it consists of an exponential representation, higher order terms mix the contributions from the QP peak with satellite features in the spectral function. This smooths out the plateau, which is no longer separate from the QP at higher temperatures, although a weak peak is visible at 900 K, and a long tail is still present at 900 K and 1500 K. Results are similar for the VBM (Fig. 11). Unlike polar insulators, such as LiF and MgO,²⁶ or the Fröhlich model in the CE approach, there is no visible signature at $+2\omega_{LO}$.

The spectral function close to the band edges was calculated along a path between high-symmetry points in reciprocal space (as commonly done in ARPES experiments). This allows us to visualize the effects of phonons, and to compare the DM and CE approaches in reciprocal space (see Fig. 12). For both approaches we can perceive at 300 K a light red color just above the yellow band at the CBM, which represents the nonpolaronic signature. In DM, the plateau is located at $2.4\omega_{\rm LO}$ above the CBM quasi-particle peak, and as the temperature increases, the plateau shifts down about 0.32 eV, which seems constant in this energy interval (the same shift can be seen in the top plot of Fig. 10 with the shift of the plateau to lower energy from 300 K to 900 K). The QP peak shifts down from -0.179 eV to -0.367 eV when the temperature increases from 300 K to 1500 K, respectively. See Table S6.1 (ESI⁺). In the CE, the satellite band is located ω_{LO} away from the CBM QP peak at 300 K. As temperature increases the QP mixes with the nonpolaronic signature and increases the broadening. There is also a huge broadening increase at the degenerate bands close to the high-symmetry point X. The changes with temperature and absolute energy resolution mean these features should be visible experimentally, and we hope to stimulate more detailed ARPES studies on intrinsic diamond as a model for nonpolarons.

To compare more quantitatively, the DM calculated ARPES was subtracted from the CE calculations in Fig. 13. The intensity range is limited between -0.3 and 0.3 to detail the view of the nonpolaronic signatures. For positive values the CE has a higher spectral function, while the opposite occurs for negative values. One can observe a broadening or/and shift effect when transiting from DM (a thin line green or grey) to CE,

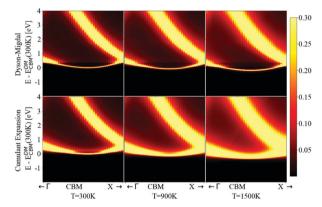
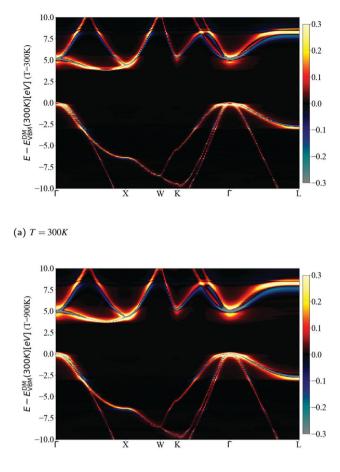


Fig. 12 Calculated ARPES spectra using DM (top) and CE (bottom) at the bottom of the conduction band. Zero was set to the CBM energy in the DM approach. The calculations were done at temperatures equal to 300, 900 and 1500 K. To be able to observe the presence of the additional signature present in the spectral function at VBM and CBM (a line just above the main QP peak on the CBM), the intensity scale of the density of states (in colors) was limited to 0.3. The plateau is visible for the different temperatures in the DM approach and the broadening changes slightly. This is opposite to CE, where we can only see the plateau at 300 K, due to mixing between the broader QP and the nonpolaronic feature.



(b) T = 900K

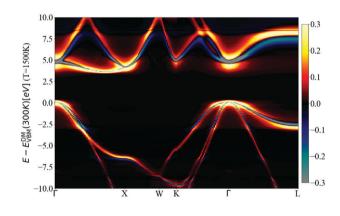




Fig. 13 Calculated spectral function difference between the CE and DM, $A^{CE}(\omega) - A^{DM}(\omega)$ at (a) 300 K, (b) 900 K, and (c) 1500 K. To highlight the nonpolaronic signature near the VBM and CBM, the intensity scale of the density of states (colormap) was limited to ± 0.3 . For positive values (red and yellow colors), the CE has more intensity than DM, and the opposite happens for the negative values (blue and grey colors). Comparatively with the narrow bands of DM, the CE approach has a much broader spread of the QP weight. Around most of the high symmetry *k*-points the broadening is almost symmetric, except *L*, where it is asymmetric, creating a tail.

which surrounds the green or grey line by a yellow or red area. In some parts of the band structure, such as the bottom of the conduction band close to the high symmetry point *L*, one can observe a tail to higher energies and an asymmetry close to the QP peak, as the weight at the QP energy using DM (blue), is spread to the tail within CE (yellow). As the temperature increases from 300 K in Fig. 13a, 900 K in Fig. 13b, to 1500 K in Fig. 13c, the broadening increases, especially for states close to the Fermi level. The shift of the QP peak from DM to CE as seen in Fig. 10 and 11 is not visible in the full band structure, as the scale of the energy in the calculated ARPES spectra is much wider than the energy shift.

Individual ARPES images for CE and DM can be found in Fig. S6.4 (ESI[†]). The dispersion of bands is noticeable and qualitatively comparable with measured ARPES images in boron-doped diamond.^{74–76} However, doping lifts the degeneracy and lowers the Fermi energy below the VBM, making it difficult to observe polaronic features, the plateau or even the QP peak at the VBM. The only observation we have found of experimental polaronic signatures in ARPES is in surface hydrogenterminated on diamond,¹² showing several satellite signatures offset by multiples of $\omega_{\rm LO}$ from the QP.

3.2 Contribution due to long-range and short-range fields

In this section, we employ the Fourier-based interpolation technique discussed in Section 2.5 to analyze the contribution given by the SR and LR quadrupolar fields to the e-ph scattering potentials. More specifically, we compare the unit cell average of the (local part) of the e-ph scattering potential⁵⁵

$$\bar{V}_{\kappa\alpha,\mathbf{q}} = \frac{1}{\Omega} \int_{\Omega} \mathrm{d}\mathbf{r} V_{\kappa\alpha,\mathbf{q}}(\mathbf{r}) e^{-i\mathbf{q}\cdot\mathbf{r}},\tag{18}$$

for several **q**-points along a high-symmetry path. Three different approaches are used to dissect the contributions to \bar{V} . In the first one, the LR contribution is given by $V_{\kappa\alpha,\mathbf{q}}^{\mathcal{L}}(\mathbf{r})$, the potential of the model in eqn (17). Dynamical quadrupoles are given by

$$Q_{\kappa\alpha}^{\beta\gamma} = (-1)^{\kappa+1} Q_{\kappa} [\varepsilon_{\beta\gamma\alpha}], \qquad (19)$$

with ε the Levi-Civita tensor, κ the atom index and Q_{κ} = 2.52. In the second approach, the SR part is obtained by Fourierinterpolating⁵⁵ the short-ranged $W_{\kappa,\alpha}(\mathbf{r}, \mathbf{R})$, in which quadrupole fields have been removed from $V_{\kappa\alpha,\mathbf{q}}(\mathbf{r})$. Finally, the total e-ph scattering potential is obtained by performing explicit DFPT calculations for all the q-points of the path. Fig. 14 compares the results obtained with these three approaches. Note, in particular, the jump discontinuity in the DFPT values for $\mathbf{q} \to \Gamma$, that is due to the LR quadrupolar potential. As discussed in ref. 5, a proper treatment of the LR quadrupole terms is necessary for an accurate interpolation of e-ph matrix elements in the long-wavelength limit. According to our results for diamond, however, an e-ph self-energy evaluated with e-ph matrix elements obtained from the LR model potential alone (eqn (17)) cannot reproduce the results obtained with e-ph matrix elements including both the SR and the LR part. In diamond, short-range crystal fields provide the most significant contribution to the QP formation with only a small influence of

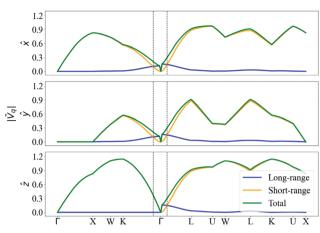


Fig. 14 Unit-cell average and absolute value of the total DFPT scattering potential (green) compared with the long-range part (blue) and the short-range contribution (orange). The three subplots corresponds to the e-ph potential generated by the displacement of a C atom along one of the three reduced directions \hat{x} , \hat{y} , \hat{z} . Both atoms present the same $|V_{\mathbf{q}}|$.

long-range quadrupoles. We can estimate the real space distance beyond which quadrupolar long-range fields dominante shortrange ones, by determining the point in the reciprocal lattice around Γ where the short-range potential becomes equal to the long-range one. By doing so, one can estimate a radius of interaction between the short-range fields and the electronic carrier. This distance is of $|\mathbf{q}| \approx 0.18$ in fractional units, represented by the two vertical dashed lines in Fig. 14. In the direct lattice it corresponds to around 18.3 Å, or around 7 unit cells in real-space, showing that diamond could be described as a medium polaron or larger. Note that we are determining an approximate interaction radius and not the size of the polaron itself. Further calculations should be carried out to estimate the polaron radius as in ref. 77 and 78.

Even if a quadrupole long-range field is present, it can be noticed from Fig. S6.3 (ESI⁺) that its importance disappears

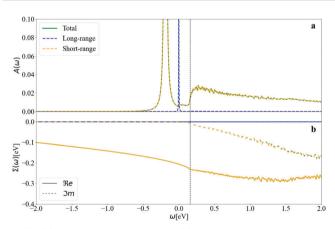


Fig. 15 (a) The spectral function and (b) the self-energy of diamond at CBM calculated at T = 300 K. The self-energy of the long-range field is negligible and leads to a single peak, at the KS energy. The short-range field calculations overlap with the total DFPT calculation, confirming it is the main contributor to the total spectra.

when calculating the scattering matrix elements, and there is thus negligible contribution from long-range fields to the nonpolaron spectral function. The maximum value of $|g_{nmk}^{jq}|$ is of 0.97, close to the effective coupling constant found experimentally in surface hydrogen-terminated on diamond of 1.1.¹²

The negligible contribution of long-range fields can also be observed in the spectral function and the self-energy in Fig. 15. The self-energy of the long-range is almost zero, meaning that there is no interaction between the electrons and the phonons that could produce long-range fields, leading to a single peak at the KS energy of the non-interacting system. The short-range field results are almost identical to those of the full DFPT potential.

4 Conclusion

First-principles e-ph calculations for diamond reveal spectroscopic signatures that originate from a quasiparticle that can be called nonpolaron: a bound electron state dressed with phonons in a non-polar system. We expect such signatures to be generic and occur in other covalent crystals.

In diamond, unexpectedly, the contribution to the formation of the nonpolaron from the long-range quadrupole fields is negligible and the binding is mainly generated by the short-range fields.

The calculations were done using both the standard Dyson-Migdal approach and the cumulant expansion. According to the CE, the signature of a nonpolaron is a plateau in the spectral function, starting one phonon energy (corresponding to the highest frequency mode) above the main QP peak. The plateau shape is at variance with the satellite peaks seen in polar systems, each separated from the QP peak or the previous replica by the LO phonon energy.

As in polar materials, we find that the QP is not properly captured in one shot NL-DM (although it might be improved by self-consistency in G). It yields an incorrect temperature dependence of the DM QP peak, which is not linear at high temperatures. The DM approximation also yields a wrong position for the plateau structure, which is located at $\pm \omega_{LO}$ relative to the KS value, independently of the position of the QP.

Calculations of the ZPR within CE and using the DM on-themass-shell approximation give similar results, and are in good agreement with experimental values of the ZPR. In addition, the cumulant expansion method delivers a physically correct nonpolaron energy, and the inclusion of the higher order terms in the e-ph interaction provides a more "realistic" view of the spectral function. The calculated ARPES spectrum of diamond around the band gap shows a broader dispersion within CE compared to Dyson–Migdal.

Finally, in order to obtain the spectral functions, selfenergies, ZPR and temperature dependence of the band gap, convergence with respect to the **q**-mesh and the broadening parameter η were carefully studied. The real part of the QP selfenergy converges quickly in opposition to the imaginary part. The narrow QP broadening at the VBM and CBM implies a very strong sensitivity of the imaginary part of the self-energy to convergence, and the imaginary η introduced in the Green's function should be chosen carefully. Our study showcases the combination of convergence techniques to reduce computational cost, in particular: the Sternheimer equation, which substitutes empty bands by a linear response equation, and the Kramers–Kronig relation, which replaces a difficult frequency integration by the use of a pre-calculated complex self-energy.

This paper motivates further research, both experimental and theoretical, on the topic of non-polar phonon modes in interacting electron–phonon systems.

Author contributions

J. C. A. prepared the manuscript, carried out calculations, prepared all graphical material and implemented the cumulant expansion, with help from M. G., J. P. N. and M. J. V. J. P. N. contributed by including supplementary information, discussions, editing and reviewing. M. G. and X. G. contributed with discussions, editing and reviewing. M. J. V. and X. G. acquired the main funding for the project. M. J. V. supervised the project and calculations, and edited and reviewed the manuscript.

Conflicts of interest

The authors have no conflicts of interest to declare.

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References

- 1 M. Meevasana, et al., New J. Phys., 2010, 12, 023004.
- 2 A. Janotti, J. B. Varley, M. Choi and C. G. Van de Walle, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 2014, **90**, 085202.
- 3 P. M. Chaikin, A. F. Garito and A. J. Heeger, *Phys. Rev. B: Solid State*, 1972, **5**, 4966.
- 4 K. P. McKenna, M. J. Wolf, A. L. Shluger, S. Lany and A. Zunger, *Phys. Rev. Lett.*, 2012, **108**, 116403.
- 5 G. Brunin, H. P.-C. Miranda, M. Giantomassi, M. Royo, M. Stengel, M. J. Verstraete, X. Gonze, G.-M. Rignanese and G. Hautier, *Phys. Rev. Lett.*, 2020, **125**, 136601.

- 6 J. Park, J.-J. Zhou, V. A. Jhalani, C. E. Dreyer and M. Bernardi, *Phys. Rev. Lett.*, 2020, **102**, 125203.
- 7 D. Emin, C. H. Seager and R. K. Quinn, *Phys. Rev. Lett.*, 1982, 28, 813.
- 8 D. J. Gibbons and W. E. Spear, J. Phys. Chem. Solids, 1966, 27, 1917.
- 9 S. H. Howe, P. G. Le Comber and W. E. Spear, *Solid State Commun.*, 1971, **9**, 65.
- 10 V. Vasilchenko, S. Levchenko, V. Perebeinos and A. Zhugayevych, J. Phys. Chem. Lett., 2021, 12, 4674.
- 11 D. Emin and M.-N. Bussac, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1994, **49**, 14290.
- 12 J. D. Rameau, J. Smedley, E. M. Muller, T. E. Kidd and P. D. Johnson, *Phys. Rev. Lett.*, 2011, **106**, 137602.
- 13 K. M. Shen, et al., Phys. Rev. Lett., 2004, 93, 267002.
- 14 M. Mohamed, M. M. May, M. Kanis, M. Brützam, R. Uecker, R. van de Krol, C. Janowitz and M. Mulazzi, *RSC Adv*.15606.
- 15 C. Chen, J. Avila, E. Frantzeskakis, A. Levy and M. C. Asensio, *Nat. Commun.*, 2015, 6, 8585.
- 16 M. Emori, A. Sakino, K. Ozawa and H. Sakama, Solid State Commun., 2014, 188, 15.
- 17 S. Moser, et al., Phys. Rev. Lett., 2013, 110, 196403.
- 18 M. Kang, et al., Nat. Mater., 2018, 17, 676.
- J. M. Riley, F. Caruso, C. Verdi, L. B. Duffy, M. D. Watson, L. Bawden, K. Volckaert, G. van der Laan, T. Hesjedal, M. Hoesch, F. Giustino and P. D.-C. King, *Nat. Commun.*, 2018, 9, 2305.
- 20 C. Cancellieri and V. N. Strocov, Spectroscopy of Complex Oxide Interfaces: Photoemission and Related Spectroscopies, Springer, Cham, 2018, pp. 107–151.
- 21 G. Antonius, S. Poncé, P. Boulanger, M. Côté and X. Gonze, *Phys. Rev. Lett.*, 2014, **112**, 215501.
- 22 P. B. Lautenschlager, P. Allen and M. Cardona, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1986, 33, 5501.
- 23 Z. Tong, S. Li, X. Ruan and H. Bao, *Phys. Rev. B*, 2019, 100, 144306.
- 24 J. M. Ziman, *Electrons and Phonons: The theory of transport phenomena in solids*, Oxford University Press, 1960.
- 25 F. Marsiglio and J. P. Carbotte, in *Electron-Phonon Superconductivity*, ed. K. H. Bennemann and J. B. Ketterson, Springer Berlin Heidelberg, Berlin, Heidelberg, 2008, pp. 73–162.
- 26 J. P. Nery, P. B. Allen, G. Antonius, L. Reining, A. Miglio and X. Gonze, *Phys. Rev. B*, 2018, **97**, 115145.
- 27 X. Gonze and C. Lee, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1997, **55**, 10355.
- 28 S. Baroni, S. Gironcoli and A. Dal Corso, *Rev. Mod. Phys.*, 2001, 73, 515.
- 29 R. M. Martin, L. Reining and D. M. Ceperley, *Interacting Electrons: Theory and Computational Approaches*, Cambridge University Press, 2016.
- 30 F. J. Dyson, Phys. Rev., 1949, 75, 1736.
- 31 A. B. Migdal, Zh. Eksp. Teor. Fiz., 1958, 34, 1438.
- 32 R. Kubo, J. Phys. Soc. Jpn., 1962, 17, 1100.
- 33 O. Gunnarsson, V. Meden and Schönhammer, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 1994, **50**, 10462.

- 35 F. Aryasetiawan, L. Hedin and K. Karlsson, *Phys. Rev. Lett.*, 1996, 77, 2268.
- 36 M. Guzzo, J. Kas, F. Sottile, M. Silly, F. Sirotti, J. Rehr and L. Reining, *Eur. Phys. J. B*, 2012, 85, 324.
- 37 F. Caruso and F. Giustino, Phys. Rev. B: Condens. Matter Mater. Phys., 2015, 92, 045123.
- 38 E. Antoncik, Czech. J. Phys., 1955, 5, 449.
- 39 H. Y. Fan, Phys. Rev., 1951, 82, 900.
- 40 F. Giustino, Rev. Mod. Phys., 2017, 89, 015003.
- 41 P. B. Allen and V. Heine, J. Phys. C: Solid State Phys., 1976, 9, 2305.
- 42 S. Poncé, Y. Gillet, J. Laflamme Janssen, A. Marini, M. Verstraete and X. Gonze, *J. Chem. Phys.*, 2015, 143, 102813.
- 43 G. D. Mahan, Many-Particle Physics, Springer, Boston, 2000.
- 44 G. Antonius, S. Poncé, E. Lantagne-Hurtubise, G. Auclair, X. Gonze and M. Côté, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 2015, **92**, 085137.
- 45 A. Mishchenko, N. Prokof'ev, A. Sakamoto and B. Svistunov, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 2000, **62**, 6317–6336.
- 46 B. Gumhalter, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 2005, 72, 165406.
- 47 D. C. Langreth, Phys. Rev. B: Solid State, 1970, 1, 471.
- 48 M. Guzzo, G. Lani, F. Sottile, P. Romaniello, M. Gatti, J. J. Kas, J. J. Rehr, M. G. Silly, F. Sirotti and L. Reining, *Phys. Rev. Lett.*, 2011, 107, 166401.
- 49 L. Hedin, Phys. Scr., 1980, 21, 477.
- 50 A. Eiguren and C. Ambrosch-Draxl, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 2008, **78**, 045124.
- 51 F. Giustino, M. L. Cohen and S. G. Louie, *Phys. Rev. B:* Condens. Matter Mater. Phys., 2007, **76**, 165108.
- 52 P. Vogl, Phys. Rev. B: Solid State, 1976, 13, 694.
- 53 C. Verdi and F. Giustion, Phys. Rev. Lett., 2015, 115, 176401.
- 54 J. Sjakste, N. Vast, M. Calandra and F. Mauri, *Phys. Rev. B:* Condens. Matter Mater. Phys., 2015, **92**, 054307.
- 55 G. Brunin, H. P.-C. Miranda, M. Giantomassi, M. Royo, M. Stengel, M. J. Verstraete, X. Gonze, G.-M. Rignanese and G. Hautier, *Phys. Rev. B*, 2020, **102**, 094308.

- 56 V. A. Jhalani, J.-J. Zhou, J. Park, C. E. Dreyer and M. Bernardi, *Phys. Rev. Lett.*, 2020, **125**, 136602.
- 57 M. Royo and M. Stengel, Phys. Rev. X, 2019, 9, 021050.
- 58 X. Gonze, et al., Comput. Phys. Commun., 2020, 248, 107042.
- 59 A. H. Romero, et al., J. Chem. Phys., 2020, 152, 124102.
- 60 D. M. Roessler and W. C. Walker, *Phys. Rev.*, 1967, **159**, 733.
- 61 D. Roessler and W. C. Walker, J. Phys. Chem. Solids, 1967, 28, 1507.
- 62 O. Madelung, *Semiconductors: Data Handbook*, Springer, London, 2004.
- 63 F. Karsai, M. Engel, E. Flage-Larsen and G. Kresse, *New J. Phys.*, 2018, **20**, 123008.
- 64 A. Miglio, V. Brousseau-Couture, E. Godbout, G. Antonius,
 Y.-H. Chan, S. G. Louie, M. Côté, M. Giantomassi and
 X. Gonze, *npj Comput. Mater.*, 2020, 6, 167.
- 65 V. Brousseau-Couture, E. Godbout, M. Côté and X. Gonze, *Phys. Rev. B*, submitted.
- 66 P. E. Blöchl, O. Jepsen and O. K. Andersen, *Phys. Rev. B:* Condens. Matter Mater. Phys., 1994, **49**, 16223.
- 67 X. Gonze, P. Boulanger and M. Côté, Ann. Phys., 2010, 523, 168–178.
- 68 R. Pässler, Phys. Status Solidi B, 1999, 216, 975.
- 69 C. D. Clark, P. J. Dean and P. V. Harris, *Proc. R. Soc. A*, 1964, 277, 312.
- 70 M. Cardona, Solid State Commun., 2005, 133, 3.
- 71 B. Monserrat, G. J. Conduit and R. J. Needs, *Phys. Rev. B: Condens. Matter Mater. Phys.*, 2014, **90**, 184302.
- 72 S. Zollner, M. Cardona and S. Gopalan, Phys. Rev. B: Condens. Matter Mater. Phys., 1992, 45, 3376.
- 73 D. L. Burk and S. L. Friedberg, Phys. Rev., 1958, 111, 1275.
- 74 T. Yokoya, T. Nakamura, T. Matsushita, T. Muro, Y. Takano,
 M. Nagao, T. Takenouchi and H. T.-O. Kawarada, *Nature*, 2005, 438, 647.
- 75 A. C.-t Pakpour-Tabrizi, Nanoscale Adv., 2020, 2, 1358.
- 76 H.-t Okazaki, J. Phys. Chem. Solids, 2008, 69, 2978.
- 77 F. M. Peeters and J. T. Devreese, Phys. Rev. B: Condens. Matter Mater. Phys., 1985, 31, 4890.
- 78 W. H. Sio, C. Verdi, S. Ponce and F. Giustino, *Phys. Rev. B*, 2019, **99**, 235139.